

**APPENDIX**

## APPENDIX A

### Theoretical Background

#### The Theory of Quantum Chemical Calculations: Ground State Methods

Molecular orbital theory is concerned with predicting the properties of atomic and molecular systems. It is based upon the fundamental laws of quantum mechanics and uses a variety of mathematical transformation and approximation techniques to solve the fundamental equations, in contrast to semi-empirical models.

#### 1. The Schrödinger Equation

The quantum chemical methods are based on finding solutions to the Schrödinger equation on molecular orbital theory. Quantum mechanics explains how entities like electrons have both particle-like and wave-like characteristics. The time independent Schrödinger equation for a molecule (n-electron and N-nuclei system):

$$H\psi(\vec{r}, \vec{R}) = E\psi(\vec{r}, \vec{R}) \quad (1)$$

and the Hamiltonian is (in atomic units):

$$H = T + V$$

$$H = -\sum_{i=1}^N \frac{1}{2} \nabla_i^2 - \sum_{A=1}^M \frac{1}{2M_A} \nabla_A^2 - \sum_{i=1}^N \sum_{A=1}^M \frac{Z_A}{r_{iA}} + \sum_{i=1}^N \sum_{j>i}^N \frac{1}{r_{ij}} + \sum_{A=1}^M \sum_{B>A}^M \frac{Z_A Z_B}{R_{AB}} \quad (2)$$

where T and V are the kinetic and potential energy operators, respectively, which separate out the motion of the nuclei from the motion of the electrons, equation (2) can be rewritten as

$$H = -\sum_{A=1}^M \frac{1}{2M_A} \nabla_A^2 + H_{\text{el}} \quad (3)$$

This focus on the electronic Hamiltonian,  $H_{\text{el}}$ , and try to solve the electronic Schrödinger equation in the field of the fixed nuclei. The nuclear-nuclear repulsion term (the final in equation (2)) appears as a constant in  $H_{\text{el}}$ . Further assume the wave function  $\psi(\vec{r}, \vec{R})$  to be a product of an electronic and a nuclear part:

$$\psi(\vec{r}, \vec{R}) = \psi_{\text{elec}}(\vec{r}, \vec{R}) \psi_{\text{nucl}}(\vec{R}) \quad (4)$$

The justification for this is that the electrons are much lighter than the nuclei. This is called the Born-Oppenheimer approximation. The parametric  $\vec{R}$  dependence of  $\psi_{\text{elec}}$  arises since the electron distribution depends implicitly on the particular nuclear arrangement for the system under study. The nuclear wave function,  $\psi_{\text{nucl}}$ , describes the vibrational, rotational and translational motion of the nuclei. From (1), (3) and (4) can obtain;

$$\begin{aligned} H\psi(\vec{r}, \vec{R}) &= H\psi_{\text{elec}}(\vec{r}, \vec{R})\psi_{\text{nucl}}(\vec{R}) = \left( -\sum_{A=1}^M \frac{1}{2M_A} \nabla_A^2 + H_{\text{el}} \right) \psi_{\text{elec}}(\vec{r}, \vec{R})\psi_{\text{nucl}}(\vec{R}) \\ &= \left( -\sum_{A=1}^M \frac{1}{2M_A} \nabla_A^2 + E_{\text{el}} \right) \psi_{\text{elec}}(\vec{r}, \vec{R})\psi_{\text{nucl}}(\vec{R}) = E\psi_{\text{elec}}(\vec{r}, \vec{R})\psi_{\text{nucl}}(\vec{R}) = E\psi \quad (5) \end{aligned}$$

The electronic wavefunction  $\psi_{\text{elec}}(\vec{r}, \vec{R})$  can be divided out from both sides of equation (5), provided that terms in  $\nabla^2 \psi_{\text{elec}}(\vec{r}, \vec{R})$  are small, i.e. the electronic wavefunction changes slowly upon small displacements of the nuclear positions. Thus, if we neglect the influence of the nuclear derivative on the electron wave function ( $\psi_{\text{el}}$ ) (i.e. the nuclei move slowly compared with the electrons) which can separate equation (5) into two equations, an electronic part:

$$H_{\text{el}}\Psi_{\text{el}}(\vec{r}, \vec{R}) = E_{\text{el}}(\vec{R})\Psi_{\text{el}}(\vec{r}, \vec{R}) \quad (6)$$

where

$$H_{\text{el}} = -\sum_{i=1}^N \frac{1}{2} \nabla_i^2 - \sum_{i=1}^N \sum_{A=1}^M \frac{Z_A}{r_{iA}} + \sum_{i=1}^N \sum_{j>i}^N \frac{1}{r_{ij}} + \sum_{A=1}^M \sum_{B>A}^M \frac{Z_A Z_B}{R_{AB}} \quad (7)$$

and a nuclear part:

$$H_{\text{nucl}}\Psi_{\text{nucl}}(\vec{R}) = E\Psi_{\text{nucl}}(\vec{R}) \quad (8)$$

where

$$H_{\text{nucl}} = -\sum_{A=1}^M \frac{1}{2M_A} \nabla_A^2 + E_{\text{el}}(\vec{R}) \quad (9)$$

## **2. Hartree Fock Theory**

An exact solution to the Schrödinger equation is not possible for any but the most trivial molecular systems. However, a number of simplifying assumptions and procedures do make an approximate solution possible for a large range of molecules. To simplify the treatment further, the next step is to assume that the electrons are non-interacting. This implies that (apart from the constant nuclear-nuclear repulsion term) which can rewrite the total n-electron Hamiltonian as a sum of n one-electron Hamiltonians,

$$H_{\text{el}} = \sum_{i=1}^N h(i) \quad (10)$$

$$h(i) = \left( -\frac{1}{2} \nabla_i^2 - \sum_{A=1}^M \frac{Z_A}{r_{iA}} \right) \quad (11)$$

This is clearly an oversimplification, since we have neglected the electron-electron repulsion term  $\frac{1}{r_{ij}}$ . Equation (10) defines the independent particle model.

The one-electron Hamiltonians (equation (11)) are termed core-Hamiltonians, since the only interactions included are those between the electrons and the bare nuclei. Including an average interaction term in the  $\{h(i)\}$ , these become effective one-electron Hamiltonians. As a consequence of equation (10), the total wave function can be rewritten as a product of  $n$  single-particle wave functions,

$$\psi(\vec{r}) = \phi_1(\vec{r}_1)\phi_2(\vec{r}_2)\dots\phi_n(\vec{r}_n) \quad (12)$$

or, take the electron spin into account,

$$\psi = \chi_1(\vec{x}_1)\chi_2(\vec{x}_2)\dots\chi_n(\vec{x}_n) \quad (13)$$

The spin orbitals  $\{\chi_i(\vec{x}_i)\}$  are the products of the spatial orbitals  $\phi_i(\vec{r}_i)$  and the spin functions ( $\alpha(\omega)$  and  $\beta(\omega)$ );  $\vec{x}_i$  denotes both the space and spin coordinates of electron  $i$ . The total independent particle spin-orbital wave function (equation (13)) is called a Hartree-product. This is an eigenfunction of the  $n$ -electron model Hamiltonian defined in equation (10), and the corresponding eigenvalue is a sum of the single-particle spin-orbital energies,

$$E_{el} = \sum_{i=1}^M \epsilon_i \quad (14)$$

A further requirement on the state wave function (13) is that it must be anti-symmetric with respect to the interchange of coordinate  $r$  (both space and spin) of any two electrons,

$$|\Psi(\bar{x}_1, \bar{x}_2, \dots, \bar{x}_n)|^2 = |\Psi(\bar{x}_2, \bar{x}_1, \dots, \bar{x}_n)|^2 \quad (15)$$

$$\Psi(\bar{x}_1, \bar{x}_2, \dots, \bar{x}_n) = \pm \Psi(\bar{x}_2, \bar{x}_1, \dots, \bar{x}_n) \quad (16)$$

It is also possible to write equation (16) in terms of a  $n \times n$  determinant, a Slater determinant, which has the same antisymmetric properties:

$$\Psi = (n!)^{-1/2} \begin{vmatrix} \chi_1(\bar{x}_1) & \chi_2(\bar{x}_1) & \cdots & \chi_n(\bar{x}_1) \\ \chi_1(\bar{x}_2) & \chi_2(\bar{x}_2) & & \\ \cdots & & \cdots & \\ \chi_1(\bar{x}_n) & \chi_2(\bar{x}_n) & \cdots & \chi_n(\bar{x}_n) \end{vmatrix} \quad (17)$$

Which commonly is written like:

$$|\Psi\rangle = (n!)^{-1/2} |\chi_1(\bar{x}_1), \chi_2(\bar{x}_2), \dots, \chi_n(\bar{x}_n)\rangle \quad (18)$$

It can easily be verified that the Slater determinant obeys the Pauli principle, as the determinant then becomes zero. The pre-factor  $(n!)^{-1/2}$  is a normalisation constant, and the  $\{\chi_i\}$  are assumed orthonormal. By antisymmetrizing the Hartree-product (13) in the form of a Slater determinant (17), that the probability of finding any two electrons at the same point in space (i.e.  $\bar{x}_1 = \bar{x}_2$ ) is zero.

Through the wave functions, the effective potential is generated. This potential allows to refine wave functions, from which a new potential is obtained. The procedure is repeated until a stable, self-consistent solution is reached. Due to the iterative procedure, the initial guess of the wave function can of course be chosen ad hoc. However, the better the initial guess is, the easier it is to reach a stable solution to the eigenvalue problems in a relatively short computational time, is provided by the variation principle. This can be stated in the following way: Given any approximate wave function, satisfying the correct boundary conditions, the expectation value of the

energy obtained by this wave function never lies below the exact energy of the ground state. Expressed in mathematical terms:

$$E_e = \frac{\langle \psi | H_e | \psi \rangle}{\langle \psi | \psi \rangle} \geq E_{\text{exact}} \quad (19)$$

A conceptually appealing model for the (trial) wave function of our molecular system is to regard it as being constructed from molecular orbitals (MO). This description is analogous to the model used for the atomic orbitals (AO). The MO's, the elements of the wave function determinant, are in turn thought of as being constructed by a Linear Combination of Atomic Orbitals (LCAO),

$$\psi_i^{\text{MO}} = \sum_{\mu} c_{\mu i} \phi_{\mu}^{\text{AO}} \quad (20)$$

The variational principle leads to following equations describing the molecular orbital expansion coefficients,  $c_{\nu i}$ , derived by Roothaan and by Hall:

$$\sum_{\nu=1}^N (F_{\mu\nu} - \epsilon_i S_{\mu\nu}) c_{\nu i} = 0 \quad \mu = 1, 2, \dots, N \quad (21)$$

Equation 21 can be rewritten in matrix form:

$$FC = SC\epsilon \quad (22)$$

with

$$F_{\mu\nu} = H_{\mu\nu}^{\text{core}} + \sum_{\lambda\sigma} P_{\lambda\sigma} \left[ (\mu\nu|\lambda\sigma) - \frac{1}{2} (\mu\lambda|\nu\sigma) \right] \quad (23)$$

$$F_{\mu\nu} = H_{\mu\nu}^{\text{core}} + G_{\mu\nu} \quad (24)$$

where  $H_{\mu\nu}^{\text{core}}$ , core-Hamiltonian matrix, defined as

$$H_{\mu\nu}^{\text{core}} = \int dr_1 \phi_{\mu}^*(1) h(1) \phi_{\nu}(1) \quad (25)$$

The matrix P is the density matrix or charge- and bond-order matrix,

$$P_{\mu\nu} = 2 \sum_a^{N/2} C_{\mu a} C_{\nu a}^* \quad (26)$$

The matrix S is the overlap matrix, indicating the overlap between orbitals.

$$S_{\mu\nu} = \int dr_1 \phi_{\mu}^*(1) \phi_{\nu}(1) \quad (27)$$

The term  $(\mu\nu|\lambda\sigma)$  in Equation 23 signified the two-electron repulsion integrals, defined as

$$(\mu\nu|\lambda\sigma) = \int dr_1 dr_2 \phi_{\mu}^*(1) \phi_{\nu}(1) r_{12}^{-1} \phi_{\lambda}^*(2) \phi_{\sigma}(2) \quad (28)$$

The (initial) wave function is used to generate an effective potential, which apply this potential in order to refine the coefficient matrix. The modified MO's form the new input in the Roothaan or Pople-Nesbet (1986) equations, and a new potential is generated. The iterative procedure is repeated until convergence is reached, i.e. when the changes in energy and/or charge density in two subsequent iterations are below a pre-set threshold value.

Before a more technical description of the SCF-procedure is presented, first need to define a new transformation matrix X, used for orthogonalisation of the basis set. This orthogonalisation can be either symmetric or canonical. A symmetric orthogonalisation implies that X is formed through the relation

$$X = S^{-1/2} = US^{-1/2}U^T \quad (29)$$

where  $S$  is the overlap matrix,  $U$  is a unitary matrix which diagonalizes  $S$ , and the diagonal matrix of the eigenvalues of  $S$  is given by the relations. In the canonical orthogonalisation procedure,  $X$  is instead given by

$$X = US^{-1/2} \quad (30)$$

Consider a new coefficient matrix  $C'$  related to the old coefficient matrix  $C$  by

$$C' = X^{-1}C, \quad C = XC' \quad (31)$$

where assumed that  $X$  possesses an inverse. Substituting  $C = XC'$  into the Roothaan equations gives

$$FXC' = SXC'\epsilon \quad (32)$$

Multiplying on the left by  $X^T$  gives

$$(X^T FX)C' = (X^T SX)C'\epsilon \quad (33)$$

if define a new matrix  $F^T$  by

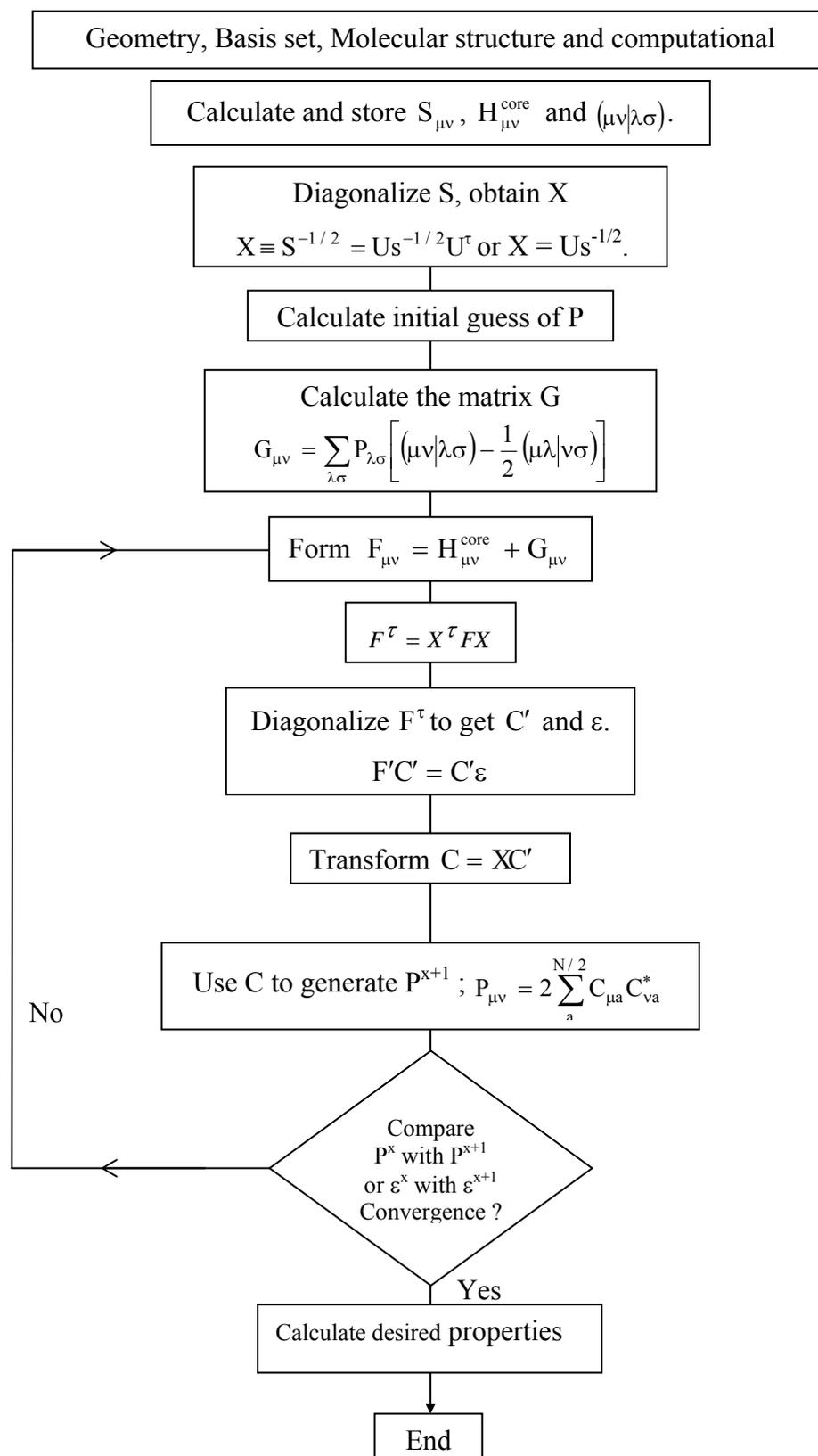
$$F^T = X^T FX \quad (34)$$

and use (27), then

$$F^T C' = C'\epsilon \quad (35)$$

The SCF procedure, outlined in Appendix figure A1, is as follows

1. Specify a molecule (a set of nuclear coordinates  $\{\mathbf{R}_A\}$ , atomic numbers  $\{Z_A\}$ , and number of electron  $N$ ) and a basis set  $\{\phi_\mu\}$ .
2. Calculate all required molecular integrals,  $S_{\mu\nu}$ ,  $H_{\mu\nu}^{\text{core}}$  and  $(\mu\nu|\lambda\sigma)$ .
3. Diagonalize the overlap matrix  $S$  and obtain a transformation matrix  $X$  from either equation  $X \equiv S^{-1/2} = US^{-1/2}U^\tau$  or  $X = US^{-1/2}$ .
4. Obtain a guess at the density matrix  $P$ .
5. Calculate the matrix  $G$  of equation  $F_{\mu\nu} = H_{\mu\nu}^{\text{core}} + G_{\mu\nu}$  from the density matrix  $P$  and the two-electron integral  $(\mu\nu|\lambda\sigma)$ .
6. Add  $G$  to the core-Hamiltonian to obtain the Fock matrix  $F = H^{\text{core}} + G$ .
7. Calculate the transformed Fock matrix  $F^\tau = X^\tau F X$ .
8. Diagonalize  $F^\tau$  to obtain  $C'$  and  $\epsilon$ .
9. Calculate  $C = X C'$ .
10. Form a new density matrix  $P$  from  $C$  using  $P_{\mu\nu} = 2 \sum_a^{N/2} C_{\mu a} C_{\nu a}^*$ .
11. Determine whether the procedure has converged, i.e. determine whether the new density matrix of step (10) is the same as the previous density matrix within a specified criterion. If the procedure has not converged, return to step (5) with the new density matrix.
12. If the procedure has converged, then use the resultant solution, represented by  $C, P, F$ , etc., to calculate expectation values and other quantities of interest.



**Appendix Figure A1** Schematic view of a Hartree-Fock self consistent field calculation.

### 3. Basis Set

The basis set most commonly used in quantum mechanical calculations are composed of atomic functions. The next approximation involves expressing the molecular orbitals as linear combinations of a pre-defined set of one-electron functions known as basis function. An individual molecular orbitals is defined as:

$$\phi_i = \sum_{\mu=1}^N c_{\mu i} \chi_{\mu} \quad (36)$$

where the coefficients  $c_{\mu i}$  are known as molecular orbital expansion coefficients. The basis function  $\chi_1 \dots \chi_N$  are also chosen to be normalized. Gaussian-type atomic functions were used as basis functions. Gaussian functions have the general form

$$g(\alpha, \vec{r}) = c x^n y^m z^l e^{-\alpha r^2} \quad (37)$$

where  $\vec{r}$  is of course composed of x, y, and z.  $\alpha$  is a constant determining the size (radical extent) of the function. In Gaussian function,  $e^{-\alpha r^2}$  is multiplied by powers (possibly 0) of x, y, and z and a constant for normalization, so that:

$$\int_{\text{allspace}} g^2 = 1 \quad (38)$$

Thus, c depends on  $\alpha$ , l, m, and n.

Here are three representative Gaussian functional (s,  $p_y$  and  $d_{xy}$  types, respectively):

$$g_s(\alpha, \vec{r}) = \left( \frac{2\alpha}{\pi} \right)^{3/4} e^{-\alpha r^2}$$

$$\begin{aligned}
 g_y(\alpha, \vec{r}) &= \left( \frac{128 \alpha^5}{\pi^3} \right)^{1/4} y e^{-\alpha r^2} \\
 g_{xy}(\alpha, \vec{r}) &= \left( \frac{2048 \alpha^7}{\pi^3} \right)^{1/4} x y e^{-\alpha r^2}
 \end{aligned}
 \tag{39}$$

Linear combinations of primitive gaussians like these are used to form the actual basis functions; the latter are called contracted Gaussians and have the form

$$\chi_\mu = \sum_p d_{\mu p} g_p
 \tag{40}$$

where the  $d_{\mu p}$ 's are fixed constants within a given basis set. Note that contracted functions are also normalized in common practice. A few commonly used basis sets are listed as following.

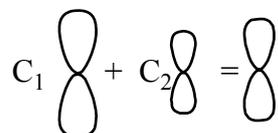
**Minimal Basis Sets:** Minimal basis sets contain the minimum number of basis functions needed for each atom, as in these examples:

H: 1s

C: 1s, 2s, 2p<sub>x</sub>, 2p<sub>y</sub>, 2p<sub>z</sub>

Minimal basis sets use fixed-size atomic-type orbitals. The STO-3G basis set is a minimal basis set (although it is not the smallest possible basis set). It used three gaussian primitives per basis function, which accounts for the "3G" in its name. "STO" stands for "Slater-type orbitals," and the STO-3G basis set approximates Slater orbitals with gaussian functions.

### Split Valence Basis Sets



The first way that a basis set can be made larger is to increase the number of basis functions per atom. Split valence basis sets, such as 3-21G and 6-31G, have two (or more) sized of basis function for each valence orbital. For example, hydrogen and carbon are represented as:

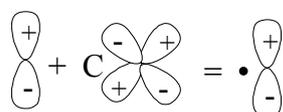
H: 1s, 1s'

C: 1s, 2s, 2s', 2p<sub>x</sub>, 2p<sub>y</sub>, 2p<sub>z</sub>, 2p'<sub>x</sub>, 2p'<sub>y</sub>, 2p'<sub>z</sub>

where the primed and unprimed orbitals differ in size.

The double zeta basis sets, such as the Dunning-Huzinaga basis set (D95), form all molecular orbitals from linear combinations of two sized of functions for each atomic orbital. Similarly, triple split valence basis sets, like 6-311G, use three sizes of contracted functions for each orbital-type.

### Polarized Basis Sets



Split valence basis sets allow orbitals to change size, but not to change shape. Polarized basis sets remove this limitation by adding orbitals with angular momentum beyond what is required for the ground state to the description of each atom. For example, polarized basis sets add d functions to carbon atoms and f functions to transition metals, and some of them add p functions to hydrogen atoms.

So far, the only polarized basis set 6-31G(d) is used. Its name indicates that it is the 6-31G basis set with d functions added to heavy atoms. This basis set is becoming very common for calculations involving up to medium-sized systems. This basis set is also known as 6-31G\*. Another popular polarized basis set is 6-31G(d,p), also known as 6-31G\*\*, which adds p functions to hydrogen atoms in addition to the d functions on heavy atoms.

### Diffuse Functions

$$C_1 \text{ (small circle)} + C_2 \text{ (large circle)} = \text{ (medium circle)}$$

Diffuse functions are large-size versions of s- and p- type functions (as opposed to the standard valence-size functions) which allow orbitals to occupy a larger region of space. Basis sets with diffuse functions are important for systems where electrons are relatively far from the nucleus: molecules with lone pairs, anions and other systems with significant negative charge, systems in their excited states, systems with low ionization potentials, descriptions of absolute acidities. The 6-31+G(d) basis set is the 6-31G(d) basis set with diffuse functions added to heavy atoms. The double plus version, 6-31++G(d), adds diffuse functions to the hydrogen atoms as well. Diffuse functions on hydrogen atoms seldom make a significant difference in accuracy.

### High Angular Momentum Basis Sets

Even larger basis sets are now practical for many systems. Such basis sets add multiple polarization functions per atom to triple zeta basis set. For example, the 6-31G(2d) basis set adds two d functions per heavy atom instead of just one, while the 6-311++G(3df,3pd) basis set contains three sets of valence region functions, diffuse functions on both heavy atoms and hydrogens, and multiple polarization functions: 3 d functions and 1 f function on heavy atoms and 3 p functions and 1 d function on

hydrogen atoms. Such basis sets are useful for describing the interactions between electrons in electron correlation methods.

#### 4. Semiempirical Calculations

Because both time and storage requirements of an *ab initio* Hartree-Fock calculation increase as the fourth power of the number of basis functions, calculations on large molecules even with the smallest basis sets are apt to be prohibitive. In such situations, the NDDO (neglect of diatomic differential overlap) formalism affords practical methods for calculating the electronic structure of large systems. Here, only one- and two-centre, two-electron integrals are considered, and the Hartree-Fock matrix, consists only of elements for which basis functions  $\mu$  and  $\nu$  are on the same atom, and basis functions  $\lambda$  and  $\sigma$  are on another atom. The individual terms are defined below (the sum  $\alpha$  is over all other atoms).

$$\begin{aligned}
 F_{\mu\mu} &= H_{\mu\mu}^{\text{core}} + \sum_{\nu} P_{\nu\nu} [\langle \mu\mu | \nu\nu \rangle - \langle \mu\nu | \mu\nu \rangle] + \sum_{\delta} \sum_{\lambda} \sum_{\sigma} P_{\lambda\sigma} \langle \mu\mu | \nu\nu \rangle \\
 F_{\mu\nu} &= H_{\mu\nu}^{\text{core}} + P_{\mu\nu} [3\langle \mu\nu | \mu\nu \rangle - \langle \mu\mu | \nu\nu \rangle] + \sum_{\delta} \sum_{\lambda} \sum_{\sigma} P_{\lambda\sigma} \langle \mu\nu | \lambda\sigma \rangle \\
 F_{\mu\lambda} &= \beta_{\mu\lambda} - \frac{1}{2} \sum_{\nu} \sum_{\sigma} P_{\lambda\sigma} \langle \mu\nu | \lambda\sigma \rangle
 \end{aligned} \tag{41}$$

The elimination of three- and four-centre integrals greatly reduces the time and storage requirements for an NDDO calculation (which now increase as the square of the number of atoms) relative to that for a full Hartree-Fock treatment.

Three levels of NDDO theory are included in *SPARTAN'S SEMI EMPIRICAL* module: MNDO Modified Neglect of Diatomic Overlap

AM1 Austin Method 1

PM3 MNDO Parameterization Method 3

In all of these formalisms, only the valence electrons are considered. The one-electron terms are given by,

$$H_{\mu\nu}^{\text{core}} = U_{\mu\nu} - Z_A \sum_{B \neq A} Z_B \langle \mu\nu | \delta\delta \rangle \quad (42)$$

Here,  $\mu$  and  $\nu$  are located on atom A and the summation is over all other atoms.  $U_{\nu\nu}$  is related to the binding energy of an electron in atomic orbital  $\nu$ , and is determined from spectroscopic data.  $U_{\nu\mu}$  is set to zero for  $\nu \neq \mu$ . The second term in equation 42 represents the attraction on an electron on atom A from the nuclear framework. The two center integral involves only the s function on atom B.  $Z_A$  is the charge of atom A without its valence electrons.

All one-centre, two-electron integrals  $(\nu\nu|\mu\mu)$  and  $(\nu\mu|\nu\mu)$  are fitted to spectroscopic data. The two-centre, two-electron repulsion integrals  $(\nu\mu|\lambda\sigma)$  are approximated by classical multipole-multipole charge interactions between atoms A and B. The multipole charge separations within an atom are treated as adjustable parameters, i.e. optimized to fit the experimentally derived one-centre integrals.

The  $\beta_{\mu\lambda}$  terms appearing in the Fock matrix (equation 41) are the one-electron, two-centre core resonance integrals and are approximated as,

$$\beta_{\mu\lambda} = \frac{\beta_{\mu} + \beta_{\lambda}}{2} S_{\mu\lambda} \quad (43)$$

where  $S_{\mu\lambda}$  is the overlap integral between Slater orbitals  $\mu$  and  $\lambda$ , and  $\beta_{\mu}$  and  $\beta_{\lambda}$  are adjustable parameters optimized using experimental thermo chemical data for simple molecules. Because all of the adjustable parameters are rooted in experimental data, these methods are known as semi-empirical. As in *ab initio* Hartree-Fock calculations, an SCF procedure is used to converge on a density matrix, and the electronic energy.

The three methods differ only in the core-repulsion terms (they also differ in the detailed parameterization). Core repulsion includes nuclear repulsion and non-valence electron-electron repulsion, which are not explicitly considered in the calculation of the electronic energy. In the MNDO model, the core repulsion energy is given by,

$$E^{\text{CR}} = \sum_{A \neq B} \sum_{B \neq A} Z_A Z_B \langle \delta(A) | \delta(B) \rangle (e^{-\sigma_A R_{AB}} + e^{-\sigma_B R_{AB}}) \quad (44)$$

where  $R_{AB}$  is the internuclear distance and  $\sigma_A$  and  $\sigma_B$  are adjustable parameters fit to give the correct empirical behavior. Details are provided in the original papers. MNDO tends to overestimate core repulsion between two atoms at van der Waals distances. For this reason, the AM1 model was developed.

In AM1 a sum of Gaussians is employed to better represent the core repulsion behaviour at van der Waals distances. PM3 uses a similar core repulsion function, but differs in the parameterisation procedure.

One advantage of methods parameterised using experimental data is their implicit inclusion of electron correlation effects. However, dependence on experimental data means that semi-empirical methods would not be expected to perform well on unusual types of molecules for which no data are available from which to construct parameters.

## 5. Density Functional Theory

Methods that are rooted in the so-called density functional theory are currently regarded as very promising since are able to include a large amount of correlation effects in a formalism that essentially requires very similar computational resources as the Hartree-Fock procedure. In fact the algorithms of the approach, in which the electron density is described in terms of one-electron basis functions, are very similar

to the single-determinant HF algorithm. This property has helped to establish density functional methods as a standard tool for chemistry and physics.

While the concept of expressing part or all of the molecular energy as a functional of the electron density goes back to the early days of quantum theory, Density Functional Theory (DFT) was put on a rigorous theoretical foundation by the Hohenberg-Kohn theorem. It states that there exists unique density  $\rho$  that yields the exact ground energy of system. The subsequent work of Kohn and Sham laid the basis for practical computational applications of the DFT to real systems. The basis of their formalism are the so-called Kohn-Sham equations.

$$H\Psi_i = E_i\Psi_i \quad (45)$$

in which the Hamiltonian  $H$  is defined as

$$H = \left( -\frac{1}{2}\nabla^2 + V_{KS} \right) \quad (46)$$

where  $V_{KS}$  is a local potential defined such that the total density of the non-interacting system

$$\rho = \sum_i |\Psi_i|^2 \quad (47)$$

is the same as the density of the “real” system.  $V_{KS}$  has the three components  $V_{ext}$ ,  $V_C$  and  $V_{XC}$  containing the nuclear and external, Coulomb potential of the electrons and the exchange-correlation interactions.

$$E_{KS} = V_{ext} + \sum_{\mu\nu} P_{\mu\nu} H_{\mu\nu} + \frac{1}{2} \sum_{\mu\nu\lambda\sigma} P_{\mu\nu} P_{\lambda\sigma} \mathfrak{S} + E_x(P) + E_C(P) \quad (48)$$

In most cases the expressions for  $E_C$  and  $E_X$  cannot be computed analytically and must be obtained by numerical methods. The key difference between the Hartree-Fock and Kohn-Sham approaches to the SCF methods is the term  $E_{XC}$ , which was mostly omitted in above discussion. In HF theory, this  $E_{XC}$  is written as

$$E_{XC}^{HF} = \frac{1}{2} \sum_{\mu\nu\lambda\sigma} (P_{\mu\nu}^\alpha P_{\lambda\sigma}^\alpha P_{\mu\nu}^\beta P_{\lambda\sigma}^\beta) (\mu\nu / \lambda\sigma) \quad (49)$$

while the KS theory introduces a functional

$$E_{XC}^{HF} = \int f(\rho^\alpha, \rho^\beta, \gamma_{\alpha\alpha}, \gamma_{\alpha\beta}, \gamma_{\beta\beta}) d\mathbf{r} \quad (50)$$

for the description. The density gradient invariants  $\gamma$  ( $\gamma_{xy} = \nabla\rho_x \nabla\rho_y$ )

### Density Functionals

It is often customary to make a partition of the density functional into an exchange and correlation part for the separation of

$$E_{XC}(\rho) = E_X(\rho) + E_C(\rho) \quad (51)$$

Although distinction between exchange and correlation contributions is somewhat artificial in the context of DFT, the above separation considerably simplifies the discussion. It should, however, be explicitly noted that the definition of  $E_C$  does not correspond to the *ab initio*  $E_C$  since correlation has, by definition, meaning only in a mean field approximation and DFT is not using such an approximation. The exchange part, on the other hand, follows closely the HF definition of exchange, does however not necessarily reproduce the exact exchange.

## Exchange

The exchange energy of a uniform spin-polarized gas of spin density  $\rho_\sigma$  is

$$E_X^S = -\sum \int \rho^\sigma(\mathbf{r}) f_X^S(\rho_\sigma(\mathbf{r})) d\mathbf{r} \quad (52)$$

with  $f_X^S(\rho_\sigma(\mathbf{r})) = \alpha_X [\rho_\sigma(\mathbf{r})]^{1/3}$  and  $\alpha_X = \frac{3}{2} \left( \frac{3}{4\pi} \right)^{1/3}$ . The exchange expression is sometimes labeled *Slater exchange*, thus the superscript *S*. This exchange expression serves as a base for other functional, which can be conveniently expressed in terms of their enhancement factor  $F_X$  over the exchange of the uniform electron gas

$$E_X = -\sum_\sigma \int \rho_\sigma(\mathbf{r}) f_X^S(\rho_\sigma(\mathbf{r})) F_X(\rho_\sigma)(\gamma_{\sigma\sigma}(\mathbf{r})) d\mathbf{r} \quad (53)$$

For instance, the exchange functional proposed by Perdew and Wang uses the following factor:

$$F_X^{PW}(s) = \left[ 1 + 0.0864 \frac{s^2}{m} + bs^4 + cs^6 \right]^m \quad (54)$$

with  $m = 1/15$ ,  $b = 14$ ,  $c = 0.2$  and  $s = (24\pi^2)^{-1/3} \sqrt{(\gamma_{\sigma\sigma} / \rho^{4/3})}$ .  $\gamma_{\sigma\sigma}$  here is again the squared density gradient  $\nabla_\omega^2$ . One of the most used exchange functionals is that of Becke 1988, which is often labeled B88 or simply B.

$$F_X^B = -\beta \rho^{1/3} \frac{x^2}{1 + 6\beta \sinh^{-1}(x)} \quad (55)$$

which uses the values  $x = \sqrt{(\gamma_{\sigma\sigma} / \rho_\sigma^{4/3})}$  and  $\beta = 0.0042$  in order to maintain correct boundary conditions. In a different approach, Perdew and Wang proposed an exchange formula that is designed from purely first principles.

$$F_X^{\text{PW91}} = \frac{1 + (a_1 s)(a_2 s) \sinh^{-1} + (a_3 + a_4 \exp(-100s^2))s^2}{1 + (a_1 s)(a_2 s) \sinh^{-1} + a_5 s^4} \quad (56)$$

where  $a_1 = 0.19645$ ,  $a_2 = 7.7956$ ,  $a_3 = 0.2743$ ,  $a_4 = -0.1508$ ,  $a_5 = 0.004$  and  $s$  the same as in eq. 54.

In practice, the three above exchange functionals are very similar, and are in fact based on minor corrections to the previous ones. Therefore they can be expected to produce very similar results. The enhancement over the simple electron gas, however, are significant enough and usually constitute a major improvement.

### Correlation

While it is possible to obtain  $E_C$  by some numerical methods from  $E_{XC}$  and the already known EX (cf. eq. 51) for the uniform electron gas, it is much more common to use separate correlation functionals. Distinction is made between local and gradient corrected functionals, referring to absence or presence of first order terms of the density  $\rho_\sigma$ . The local functional proposed by Vosko, Wilk and Nusair (VWN) was obtained using Pad's approximated interpolations of Ceperley and Alder results of their accurate quantum Monte Carlo calculations for the homogeneous electron gas. The functional is,

$$E_C^{\text{vwn}} = \frac{A}{2} \left[ \ln \frac{x^2}{X(x)} + \frac{2b}{Q} \tan^{-1} \frac{Q}{2x+b} - \frac{bx_0}{X(x_0)} \left( \ln \frac{(x-x_0)^2}{X(x)} + \frac{2(b+2x_0)}{Q} \tan^{-1} \frac{Q}{(2x+b)} \right) \right] \quad (57)$$

where the functions  $x = r_b^{1/2}$ ,  $X(x) = x^2 + bx + c$  and  $Q = (4c - b^2)^{1/2}$  and the constants are  $A = 0.0621814$ ,  $x_0 = -0.409286$ ,  $b = 13.0720$  and  $c = 42.7189$ .  $r_b$  represents the Wigner-radius and is defined by  $1/\rho = \frac{4\pi}{3}(r_b)^3$ . Together with the exchange expression from eq. 52 this constitutes what is often called the local density

approximation (LDA) or local spin density approximation (LSDA) when spin is considered (Gies and Gerhardt, 1987).

Due to the experiences with the LDA and as a consequence of some of its shortcomings, recent developments have resulted in a number of gradient corrections to local functionals like the aforementioned VWN or a completely new class of gradient corrected functionals.

Another frequently used functional has been published by Lee, Yang and Parr. It replaces both the local and the gradient part of the LDA correlation functional.

$$E_C^{\text{LYP}} = -a \frac{1}{1+d\rho^{-1/3}} \left\{ \rho + b\rho^{-2/3} \left[ C_F \rho^{-5/3} - 2t_w + \frac{1}{9} \left( t_w + \frac{1}{2} \nabla^2 \rho \right) \right] e^{-2c\rho^{-1/3}} \right\} \quad (58)$$

where  $t_w = \frac{1}{8} \left( \frac{|\nabla \rho|^2}{\rho} - \nabla^2 \rho \right)$  and  $C_F = \frac{3}{10} (3\pi^2)^{2/3}$ ,  $a = 0.04918$ ,  $b = 0.132$ ,  $c = 0.2533$

and  $d = 0.349$

### Hybrid Functionals

More recently, following an approach proposed by Becke, the combination of DFT functionals with *ab initio* formulations led to a class of expressions which are essentially a mixture of both DFT and HF contributions with fitted coefficients for each contribution. The aim of this approach is to provide expressions that include the full exchange contribution and avoid side-effects arising from a complete replacement of the DFT exchange expression by the HF one. As an example, the B3LYP functional looks like this:

$$E_{\text{XC}}^{\text{B3LYP}} = a_{\text{x0}} E_{\text{X}}^{\text{S}} + (1 - a_{\text{x0}}) E_{\text{X}}^{\text{HF}} + a_{\text{x1}} \Delta E_{\text{X}}^{\text{B}} + E_{\text{C}}^{\text{VWN}} + a_{\text{c}} \Delta E_{\text{C}}^{\text{LYP}} \quad (59)$$

with  $a_{x_0} = 0.80$ ,  $a_{x_1} = 0.72$  and  $a_c = 0.81$ , which are values fitted for a selected set of molecules to reproduce the heat of formation. The term  $E_X^{\text{HF}}$  is calculated using the Kohn-Sham orbitals in the manner of the HF procedure by computing the exchange integrals  $(\mu\nu / \nu\mu)$ . The B3LYP functional often uses  $\Delta E_C^{\text{3LYP}} = E_C^{\text{LYP}} - E_C^{\text{VWN}}$ .

## **The Theory of Quantum Chemical Calculations: Excited State Methods**

### **1. Excited Slater Determinants**

The Hartree-Fock method generates solutions to the Schrodinger equation where the real electron-electron interaction is replaced by an average interaction. In the sufficiently large basis, the HF wave function is able to account for  $\sim 99\%$  of the total energy, but the remaining  $\sim 1\%$  is often very important for describing chemical phenomena. It is therefore clear that in order to improve on HF results, the starting point must be a trial wave function which contains more than one Slater Determinant (SD)  $\Phi$ . Then electron correlation methods normally use the HF wave function as a starting point for improvements.

A generic multi-determinant trial wave function can be written as

$$\Psi = a_0\Phi_{\text{HF}} + \sum_{i=1} a_i \Phi_i \quad (1)$$

where  $a_0$  usually is close to 1. Electron correlation methods differ in how they calculate the coefficients in front of the other determinants,  $a_0$  being determined by the normalization condition.

How are the additional determinants beyond the HF constructed? With  $N$  electrons and  $M$  basis functions, solution of the Roothaan-Hall equations for the RHF case will yield  $N/2$  occupied MOs and  $M - N/2$  unoccupied (virtual) MOs. A Slater determinant is formed by  $N/2$  spatial MOs multiplied by two spin functions to yield  $N$  spinorbitals. By replacing MOs which are occupied in the HF determinant by MOs

which are unoccupied, a whole series of determinants may be generated as shown in Figure A2. These can be denoted according to how many occupied HF MOs have been replaced by unoccupied MOs, i.e. Slater determinants which are *singly*, *doubly*, *triply*, *quadruply* etc. excited relative to the HF determinants, up to a maximum of  $N$  excited electrons. These determinants are often referred to as *Singles* (S), *Doubles* (D), *Triples* (T), *Quadruples* (Q) etc.

The total number of determinants that can be generated depends on the size of the basis set, the larger the basis, the more virtual MOs, and the more excited determinants can be constructed. If all possible determinants in a given basis set are included, all the electron correlation (in the given basis) is (or can be) recovered. Methods which include electron correlation are thus two-dimensional, the larger the one-electron expansion (basis set size) and the larger the many-electron expansion (number of determinants), the better are the results.

| HF | S-type | S-type | D-type | D-type | T-type | Q-type |
|----|--------|--------|--------|--------|--------|--------|
| —  | —      | —      | —      | —      | —      | —      |
| —  | —      | ↑      | ↓      | ↑↓     | ↑↓     | ↑↓     |
| —  | —      | —      | —      | —      | —      | ↓      |
| —  | —      | —      | —      | —      | ↑      | ↑      |
| —  | ↑      | —      | ↑      | —      | —      | —      |
| ↑↓ | ↓      | ↑↓     | ↓      | ↑↓     | ↓      | ↓      |
| ↑↓ | ↑↓     | ↑      | ↑      | —      | —      | —      |
| ↑↓ | ↑↓     | ↑↓     | ↑↓     | ↑↓     | ↑↓     | ↑      |

**Appendix Figure A2** Excited Slater determinants generated from a HF reference.

## 2. Molecular Orbital Theories for Excited States

### 2.1. Configuration Interaction (CI)

The term “configuration” was originally used for the interaction of just a few Slater determinants required for the proper description of some electronic states of atoms and molecules, though the concept of configuration interaction has been known as a tool for obtaining, at least in principle, the exact solution of the Schrodinger equation since the early days of quantum mechanics. The results of CI calculations, and especially the results of full configuration (FCI) calculations, have been used as benchmarks for testing the newly developed methods. FCI is still the only method which is used for the definition of the correlation energy as the difference between the exact nonrelativistic energy (represented by the FCI energy) and the Hartree-Fock (HF) energy:

$$E_{\text{corr}} = E_{\text{exact}} - E_{\text{HF}} \quad (2)$$

CI is still a standard method for calculation of excited electronic states and its conceptual simplicity makes it the method of choice for qualitative explanation of phenomena for which the single-determinant Hartree-Fock approximation is unsatisfactory because of near-degeneracies or rearrangement of electrons within partly occupied shells.

#### 2.1.1 Definition and the essence of configuration interaction

For any CI calculation, we need first an orbital set, i.e., a set of one-electron functions  $\varphi = \{\varphi_i, i = 1, 2, \dots, m\}$ , that satisfy the orthonormality

$$\langle \varphi_i | \varphi_j \rangle = \delta_{ij} \quad (3)$$

and the completeness conditions

$$\sum_i |\varphi_i\rangle \langle \varphi_i| = 1 \quad (4)$$

In the most applications these orbitals are chosen to be Hartree-Fock orbitals, though as will be noted below other types of orbitals (Huckel orbitals, natural orbitals, or even nonorthogonal orbitals) may also be used.

Next that the orbitals are in a certain order, for example in the order of their increasing energy. Then assign to each orbital  $\varphi_i$  an occupation number  $n_i$ , which may take the 0, 1, or 2. The sum of the occupation numbers must be equal to the number of electrons. A set of occupation numbers  $(n_1, n_2, \dots, n_m)$  is called a *configuration*. In the ground state configuration of a closed-shell system all the lowest orbitals are doubly occupied. For example, for a ten electron system and the orbital set with  $m = 7$ , the ground state configuration is (2222200). The configuration (2222110), corresponding to the excitation of a single electron from orbital 5 to orbital 6, is called a singly excited configuration.

For a configuration having  $n$  singly occupied orbitals, we can in general form  $2^n$  spin functions and therefore also  $2^n$  different Slater determinants. A true number of linearly independent spin functions for a given configuration, i.e., for an orbital product function, and a total spin  $S$  is given by following formula:

$$f(n, S) = \binom{n}{q} - \binom{n}{q-1} \quad (5)$$

where

$$q = \frac{1}{2}n - S \quad (6)$$

In general the selected functions are linear combinations of several Slater determinants and, in contrast to individual Slater determinants, they are eigenfunctions of the spin operators  $S_z$  and  $S^2$ . Therefore they are called spin-adapted configuration functions or, most frequently, *configuration state functions* (CSFs). The

spin adapted reference configuration is called the reference state function. Because of orthogonality of orbitals and spin functions, the CSFs are also orthogonal, and because of the completeness of the orbital set, a set of all possible CSFs forms a complete set

$$\sum_i |\Phi_i\rangle \langle \Phi_i| = 1 \quad (7)$$

Hence the exact wave function for a particular orbitals set may be expressed as

$$\Psi = \sum_i |\Phi_i\rangle \langle \Phi_i | \Psi \rangle \quad (8)$$

where the overlap integrals between CSFs and the exact wave function have the meaning of expansion coefficients.

$$\Psi = \sum_i C_i \Phi_i \quad (9)$$

Use of a complete set of CSFs in the expansion (8) is referred to as the complete or full configuration (FCI). Since FCI calculations are computationally very demanding, they are only feasible for small sets and a small number of electrons. In chemical applications of the CI method, smaller expansions are used in order to develop a computationally tractable model. Standard levels of the CI method are termed according to the extent of the CI expansion: CIS, CISD, CISDT, and CISDTQ correspond to expansions through singly, doubly, triply and quadruply excited CSFs, respectively. Those methods will be called truncated CI methods. Truncating the excitation level at 1 (CI with Singles, CIS) does not give any improvement over the HF result as all matrix elements between the HF wave function and singly excited determinants are zero. CIS is equal to HF for the ground-state energy, although higher roots from the secular equations may be used as approximations to excited states. It has already been mentioned that only doubly excited determinants have matrix elements with the HF wave function different from zero, thus the lowest CI level which gives an improvement over the HF result is that which includes only doubly excited states, yielding the CID. Although the singly excited determinants have zero

matrix elements with the HF reference, they enter the wave function indirectly as they have non-zero matrix elements with the doubly excited determinants. So that the only CI method which is generally applicable for a large variety of systems is CISD. For molecular properties, the singly excited states thus allow the CI wave function to “relax” the MOs, i.e. letting the wave function respond to the perturbation.

### **2.1.2 Size Consistency, Performance, and Accuracy**

#### **Size Consistency**

Size consistency is the property that is required for any theoretical model incorporating electron correlation to be applicable to problems of chemical reactivity. It means that the method when applied to an ensemble of isolated molecules should give results which are additive for the energy and other properties. Unfortunately, any incomplete CI does not satisfy this requirement and its applications therefore must be treated with caution if comparison is made for properties of molecule of different size. The origin of this deficiency is well understood and the terms contained in the CI energy, which do not scale linearly with the number of particles, have been identified. The most convenient tool for the examination of size inconsistency in CI is the many-body perturbation theory. It was found that, unlike in the fourth- and higher-order many-body perturbation theory, the unphysical terms in CISD are not canceled and their cancellation is only achieved if quadruply and higher excited CSFs are included in the CI expansion. Complete cancellation is secured only by full CI, though it is believed that CISDTQ properties should be very near the full CI limit for chemical systems which can be well represented by a single restricted HF wave function. For dimer  $(\text{H}_2\text{O})_2$  and the double zeta basis set the CISDTQ size-inconsistency error amounts to 0.0003 au. which is considerably less than the CISD error of 0.0047 au. In general, we must expect that CISD does not provide results with the accuracy sufficient, for example, for thermochemistry. This stimulated attempts to eliminate unphysical terms in the CISD energy in some computationally simple, though

approximate way. Several formulas were suggested for the a posteriori correction of the CISD energy. The most popular among these is the Davidson correction:

$$\Delta E = (1 - C_0^2)(E_{\text{CID}} - E_{\text{SCF}}) \quad (10)$$

where  $E_{\text{CID}}$  is the CID energy and  $C_0$  is the expansion coefficient for the ground state CSF. Originally the Davidson correction was developed for CID but it is commonly used for correction of CISD energies.

### **Performance and Accuracy**

It has been stated that CISDTQ properties are very near to the full CI data for systems which can be well represented by a single restricted HF wave function. Unfortunately, in spite of the progress made recently in CISDTQ calculation, the method cannot be advocated for general use because of the immense number of triply and quadruply excited CSFs. The method scales as  $n^{10}$ , where  $n$  is the number of active molecular orbitals, and it can be used only for the smallest molecular systems. On the other hand, the CISD method scales only as  $N^6$ ,  $N$  being the number of (atomic) basis set functions, which makes the method amenable to routine calculations. In general, the errors in CISD energy predictions (heats of reactions, activation barriers) have to be expected to be considerably larger than that corresponding to the ‘chemical accuracy’ (1 kcal/mol<sup>1</sup>), even if the Davidson correction is applied. CISD fails particularly in those cases where the size-inconsistency effect is important.

If a higher accuracy is required, it is recommended to pass to the coupled cluster (CCs) theory. Basically, CCSD is also an  $N^6$  procedure, but the test calculations showed that the timing ratio CCSD/CISD ranged from 4 to 8, though as with any comparison of two different methods, the ratio depends strongly on the efficiency of the two programs.

## 2.2 Symmetry Adapted Cluster and Symmetry Adapted Cluster Configuration (SAC/SAC-CI) Method

### 2.2.1 SAC theory for the ground state

SAC is abbreviation for the symmetry adapted cluster, the meaning of which will become clear later. It belongs to the cluster expansion approach, which was originated in the statistical theory of interacting atoms, in the theory of electron correlations in atoms and molecules.

Electron correlation is defined on the basis of the HF theory as

$$E_{corr} = E_{exact} - E_{HF} \quad (11)$$

$$\chi_{corr} = \Psi_{exact} - \Psi_{HF} \quad (12)$$

where ‘exact’ stands for the exact solution of the non-relativistic Schrodinger equation. Since the HF model is independent particle model, electron correlations represent mainly the collisions of electrons scattering into unoccupied orbitals. We introduce an excitation operator  $T_I^+$  which represents such a collision. For example, a collision of two electrons belonging to the occupied orbital  $\varphi_i$ , resulting in the scattering into the unoccupied orbital  $\varphi_a$ , is represented by the excitation operator  $T_{ii}^{aa}$ ,

$$T_{ii}^{aa} = a_{a\alpha}^+ a_{a\beta}^+ a_{i\alpha} a_{i\beta} \quad (13)$$

$$T_{ii}^{aa} |0\rangle = \|\varphi_1\alpha\varphi_1\beta\dots\varphi_a\alpha\varphi_a\beta\dots\varphi_n\alpha\varphi_n\beta\|, \text{ where } |0\rangle = \Psi_g^{HF} \quad (14)$$

Configuration interaction (CI) method is one of the most popular methods for including electron correlations. This method is based on the expansion theorem, and the correlated wave function is expressed as

$$\Psi^{\text{CI}} = B_0 |0\rangle + \sum_I B_I T_I^+ |0\rangle \quad (15)$$

where  $B_I$  are expansion coefficients. This method is simple and exact, but is usually slowly converging, especially for excited states. The dimension of the configurations easily reaches to the order of  $10^8$ , though many efficient algorithms for handling such large matrices are proposed. Further, it is difficult to extract a physical meaning from such a large number of configurations. It is also difficult to solve many lower solutions of such a large matrix, which is necessary for studying shake-up spectra, for example.

The main factor of electron correlation is collisions of two electrons. In many electron systems, however, there is a chance for three, four and more electrons to collide each other. However, the probability for four collisions actually important are the products of pair collisions occurring at different places of the molecule. This is because the fluctuation potential for the electron correlation is very short range. When we introduce a sum of the excitation operation as

$$T = \sum_I C_I T_I^+ \quad (16)$$

The wave function including higher-order collisional effects is written as

$$\Psi_g = (1 + T + 1/2 T^2 + 1/6 T^3 + \dots) |0\rangle \quad (17)$$

where the terms  $T^2$ ,  $T^3$ , etc., represent two pair collisions, three pair collisions, etc., and the factors 1/2, 1/6, etc., are due to the indistinguishability of pair collisions. Eq. (17) is more compactly written as

$$\Psi_g = \exp(T) |0\rangle \quad (18)$$

which is the cluster expansion. The suffix g again stands for the ground state. The theory based on this expansion is called coupled cluster (CC) theory. Hereafter, we

call the term,  $(1+T) |0\rangle$  as linked term and term,  $(1/2 T^2 + 1/6 T^3 + \dots) |0\rangle$  as unlinked term.

In the above formulation, we have introduced the operators  $T_i^+$  representing two electron excitations (pair collisions). However, generally speaking, this is just an example, and we may take any operators physically important. An important example is to choose  $T$  as a sum of all single excitation operators,

$$T^{(1)} = \sum_i^{occ} \sum_a^{unocc} c_i^a a_a^+ a_i \quad (19)$$

where  $i$  and  $a$  stand for the general spin orbital. Then we get Thouless theorem,

$$\Psi^{SD} = N \exp (T^{(1)}) \Phi^{SD} \quad (20)$$

where  $\Phi^{SD}$  and  $\Psi^{SD}$  are different single determinants,  $N$  a normalization constant. This theorem states that the transformation of single determinants to another one is expressed by the operator  $N \exp (T^{(1)})$ : the cluster expansion includes the *self-consistency of orbitals*.

We note that for open-shell systems, the single determinant  $\Psi^{SD}$  on the left hand side of eq. (20) is not a restricted determinant, but an unrestricted one which is not an eigenfunction of the spin-squared operator  $S^2$ . Generally, the wave function of the CC theory is not an eigenfunction of  $S^2$ , as actually reported for the CCSD wave functions for doublet radicals. In the linear expansions like CI, the solution of the secular equation is always symmetry-adapted, irrespective of the choice of the excited configurations, because the Hamiltonian is totally symmetric. However, this is not the case for the non-linear expansions like cluster expansion. Further, as explained below, coupled cluster expansion may involve a larger number of variables than that necessary for describing the state.

These difficulties do not occur when excitation operations are chosen to be symmetry adapted. Excitation operator  $S_I^+$  are defined to be symmetry adapted when the configuration  $S_I^+ |0\rangle$  is symmetry adapted. For totally symmetric singlet states, the symmetry adapted cluster (SAC) expansion are defined as

$$\Psi_g^{SAC} = \exp(S) |0\rangle \quad (21)$$

where

$$S = \sum_I C_I S_I^+ \quad (22)$$

Since  $S_I^+$  is totally symmetric, the unlinked terms of eq. (21) are also totally symmetric. For open-shell states like doublet and triplet states, we need a symmetry projector  $Q$  as

$$\Psi_g^{SAC} = Q \exp(S) |0\rangle = [1 + S + Q(1/2S^2 + 1/6S^3 + \dots)] |0\rangle \quad (23)$$

where  $|0\rangle$  is a restricted determinant and  $Q$  applies only to the unlinked terms, since the linked term is already symmetry adapted.

The SAC expansion defined by eq. (21) is thus different from the CC expansion given by eq. (18). Table 1 is summarized the differences in schematic way.

**Appendix Table A1** Schematic summary of the differences of the SAC expansion from the conventional cluster expansions<sup>a</sup>.

| Expansion            | Linked term                          | Unliked term <sup>b</sup> | Number of independent variables <sup>c</sup> | Symmetry <sup>d</sup> |
|----------------------|--------------------------------------|---------------------------|--|-----------------------|
| $\exp(T)  0\rangle$  | $T 0\rangle$                         | $T_I T_J  0\rangle$       | Larger                                       | Mixed                 |
| $Q\exp(T)  0\rangle$ | $QT 0\rangle \rightarrow S 0\rangle$ | $QT_I T_J  0\rangle$      | Larger                                       | Pure                  |
| $Q\exp(S)  0\rangle$ | $S 0\rangle$                         | $QS_I S_J  0\rangle$      | Just as required                             | Pure                  |

<sup>a</sup> The operators  $T_I$  are not symmetry-adapted, but the operators  $S_I$  are symmetry-adapted.

<sup>b</sup> Only the second-order unliked terms are given

<sup>c</sup> The number of the independent variables included in each expansion is compared with that necessary for the description of the system under consideration.

<sup>d</sup> Symmetry of the total wave function.

Now the solution of SAC theory was considered. In the SAC expansion, the unknown variables  $C_I$  are associated to the linked excitation operator  $S_I^+$ , so that we require the Schrodinger equation,  $H-E_g |\Psi_g^{SAC}\rangle = 0$ , within the space of the linked configurations as

$$\langle 0 | H-E_g |\Psi_g^{SAC}\rangle = 0 \quad (24a)$$

$$\langle 0 | S_I(H-E_g) |\Psi_g^{SAC}\rangle = 0 \quad (24b)$$

We have the same number of equations as the number of the unknown variables, This solution is called non-variational solution.

The variational solution is obtained by applying the variational principle to the SAC wave function and we obtain,

$$\langle \Psi_g^{SAC} | S_I(H-E_g) | \Psi_g^{SAC} \rangle = 0 \quad (25a)$$

$$\langle \Psi_g^{SAC} | (H-E_g) S_I | \Psi_g^{SAC} \rangle = 0 \quad (25b)$$

This equation is valid *only* for the SAC expansion, but not for the CC expansion, because of the reason summarized in Table 1. Generally, the variation solution is more difficult than the non-variational one, because former involves the integrals between the unlinked terms. However, we believe, as long as the wave function itself is accurate, the difference between the variational and non-variational solutions should be small.

It is shown that eq. (25b) is the *generalized Brillouin theorem*. In comparison with eq. (26).

$$\langle \Psi_g^{HF} | H | \phi_i^a \rangle = 0 \quad (26)$$

where  $\Psi_g^{SAC}$  corresponds to  $\Psi_g^{HF}$  and  $S_I^+ | \Psi_g^{SAC} \rangle$  does to  $\phi_i^a$ . As the Brillouin theorem is a key equation in the HF/SECI theory, the generalized Brillouin theorem given by eq. (25b) is a key equation in the theoretical framework of the SAC/SAC-CI theory. The SAC theory has the following properties. The first three are common to the CC theory.

- (1) It effectively involves higher-order effects of electron collisions. It describes dynamic correlations quite effectively.
- (2) It is size consistent or size extensive, so that it correctly describes the energy change in the dissociation process such as  $X_n \rightarrow nX$ . This property is a direct consequence of the exponential, since

$$\exp(A) \exp(B) = \exp(A+B) \quad (27)$$

when the operators A and B are commutable.

(3) It includes self-consistency. This property is best represented by the Thouless' theorem given by eq. (20). It guarantees that the cluster expansion is *independent* of a choice of the reference orbitals, when we include all the single excitation operations.

(4) The SAC theory defines not only the SAC wave function itself  $\Psi_g^{SAC}$  for the ground state, but also the excited functions which span the basis for excited states. The SAC-CI theory is based on this property. This property is probably the most important property, among others, and is valid only for the SAC theory.

### 2.2.2 SAC-CI theory for excited, ionized, and electron attached states

It was thought for a long time that the description of electron correlations in excited states is much more difficult than that in the ground state. Excited states are generally open shells and are not represented by a single Slater determinant. Many different states of many different symmetries and natures are involved in a narrow energy range, which makes it difficult to suppose a single general theory in a useful form. However, this is not the case in the SAC-CI theory. By using the SAC-CI method, we can *easily* calculate the correlated wave function of the excited, ionized and electron attached states, as explained in the following.

The electron correlations in the excited state will be described on the basis of those in the ground state. Approximately, excitations and ionizations involve only one or two electrons, and most other electrons lie essentially in the same orbitals as in the ground state. Therefore, the electron correlations in the excited state should be able to be compactly described by considering only some modifications to the ground-state electron correlations.

The excited functions  $\{\Phi_K\}$  by using the SAC wave function was written as

$$\Phi_K = P S_K^+ \Psi_g^{SAC} \quad (28)$$

where  $P$  is the operator which projects out the ground state wave function,

$$P = 1 - \left| \Psi_g^{SAC} \right\rangle \left\langle \Psi_g^{SAC} \right| \quad (29)$$

and  $\{S_K^+\}$  a set of the excitation operations involving the excitations under consideration in a orbital picture. From the generalized Brillouin theorem of the SAC theory, eq. (25b), it is easily shown that the functions  $\{\Phi_K\}$  satisfy

$$\left\langle \Phi_K \left| \Psi_g^{SAC} \right\rangle = 0, \quad \left\langle \Phi_K \left| H \right| \Psi_g^{SAC} \right\rangle = 0 \quad (30)$$

These equation shows, that *the set of the functions spans the space for the excited states*. We therefore describe the excited state by *a linear combination* of the functions  $\{\Phi_K\}$ ,

$$\Psi_e^{SAC-CI} = \sum_K d_K \Phi_K \quad (31)$$

which is the SAC-CI theory.

Obviously, the SAC-CI wave function for the excited state satisfied the correct relations with the SAC ground state,

$$\left\langle \Psi_g^{SAC} \left| \Psi_g^{SAC-CI} \right\rangle = 0, \quad \left\langle \Psi_g^{SAC} \left| H \right| \Psi_g^{SAC-CI} \right\rangle = 0 \quad (32)$$

Applying the variational principle to eq. (32) for solving the unknown variables  $\{d_K\}$ , we obtain

$$\left\langle \Phi_K \left| H - E_e \right| \Psi_g^{SAC-CI} \right\rangle = 0 \quad (33)$$

Different solutions of eq. (32), which correspond to different excited states, satisfy

$$\langle \Psi_e^{SAC-CI} | \Psi_f^{SAC-CI} \rangle = 0, \quad \langle \Psi_e^{SAC-CI} | H | \Psi_f^{SAC-CI} \rangle \quad (34)$$

Since they are the solutions of the common secular equation. Thus, the SAC-CI wave function satisfies the correct relations with the ground state and with the other excited states. This is very important when we consider the properties, like transitions and relaxations, which interconnect different states.

In the above formulation, we have considered implicitly the excited states having the same symmetry as the ground state. However, the SAC-CI theory is also valid for the excited states having different symmetries (e.g., triplet), and for the ionized and electron attached states. We generalize eq. (28) as

$$\Phi_K = P R_K^+ \Psi_g^{SAC} \quad (35)$$

where  $\{R_K^+\}$  represents a set of excitation, ionization, and/or electron attachment operators. In any cases, eqs. (29)-(34) are valid.

Though the above formulation of the SAC-CI theory is variational and non-variational formulation. The non-variational SAC-CI solution is obtained by projecting the Schrodinger equation onto the space of the linked configurations,

$$\langle 0 | R_K (H - E_e) | \Psi_g^{SAC-CI} \rangle = 0 \quad (36)$$

Referring to eq. (25), we note that in the non-variational case, the SAC and SAC-CI wave functions satisfy the common set of equations. In particular, when we consider the excited states belonging to the same symmetry as the ground state, the operator  $\{R_K^+\}$  in eq. (35) are actually  $\{S_I^+\}$ . The solutions of the Schrodinger equation belonging to different eigenvalues are orthogonal and Hamiltonian orthogonal. Therefore, we obtain eqs. (32) and (34) within the space of the linked operators under

consideration. These equations are quite important for the theoretical consistency of the different states under consideration.

Practically, the non-variational solution is easier than the variational one by the reason similar to that stated for the SAC solution, but we have to diagonalize non-symmetric matrices. When the SAC-CI program was first coded, it was no efficient method for diagonalizing non-symmetric matrices of large dimensions, Therefore, we had to prepare the algorithm of iterative diagonalizations of non-symmetric, extending the Davidson 's algorithm for symmetric matrices.

As the SAC theory is exact, the SAC-CI theory is also exact. Though the introduction of the SAC-CI theory so far given is rather formal and straightforward, it has some interesting physics. First, omitting the projector, or including the identity operator into  $\{R_K^+\}$ , then the SAC-CI wave function was written as

$$\Psi_e^{SAC-CI} = \mathfrak{R} \Psi_g^{SAC} \quad (37a)$$

where the excitatory  $\mathfrak{R}$  (a king of reaction operator) is defined as

$$\mathfrak{R} = \sum_K d_K R_K^+ \quad (37b)$$

We already know that the SAC wave function well describes the electron correlation of the ground state. The excitatory  $\mathfrak{R}$  describes the excitation starting from the electron correlation involved in the SAC ground state  $\Psi_g^{SAC}$ .

Generally, excitation is only one or two electron processes and most other electrons lie in the situation (orbitals) similar to those in the ground state. Therefore, it is clever to start from the ground-state electron correlation and describe only the modifications caused by the excitation. eqs. (37a) and (37b) just represent such an idea, which is the *transferability* of electron correlations between ground and excited states. This method is much easier than calculation all of the electron correlations of

each state from the beginning. Since the ground-state electron correlation is easier to calculate than the excited state one, when we calculate from the beginning, we first calculate it by the SAC method and then utilize it in the SAC-CI method for calculating the excited-state correlations based on its transferability. For this reason, the SAC-CI expansion is much easier and more rapidly convergent than ordinary CI.

The SAC-CI wave function is also written in the form

$$\Psi_e^{SAC-CI} = \exp \left( \sum_I C_I S_I^+ \right) \sum_K d_K R_K^+ |0\rangle \quad (38)$$

which has the structure of multi-reference CI. The configurations  $R_K^+ |0\rangle$  represent the reference configurations and the operator  $\exp \left( \sum_I C_I S_I^+ \right)$  represents the excitations from these reference configurations. In the latter, we use the coefficients  $\{C_I\}$  determined for the ground state, which is based on the transferability of electron correlations between the ground and excited states. The dimension of the SAC-CI method is the number of the reference configurations which are typically in the order of  $10^3$ - $10^5$ , not like limited to  $\sim 10$ . Therefore, in the SAC-CI method, there is almost no ambiguity in the choice of the ‘main reference’ configurations in contrast to the ordinary multireference CI method.

The SAC-CI method can be applied to various kinds of excited states by using appropriate excitation operators. It can be applied to excited states, ionized states, and electron attached states having spin multiplicities of singlet, doublet, triplet and up to septet. An important merit is that we can calculate these different electronics states in a *same accuracy*. *We can directly compare* the energies and the wave functions of different electronic states: a property quite important and useful in actual applications. *Thus using the SAC-CI method, we can study chemistry and physics involving the ground state and excited states of various spin multiplicities.*

### 2.3. Time Dependent Density Functional Theory

A DFT method that is strongly analogous to RPA is called time-dependent DFT (TDDFT). In this case, the Kohn-Sham (KS) orbital energies and various exchange integrals are used in place of matrix elements of the Hamiltonian. TDDFT is usually most successful for low-energy excitations, because the KS orbital energies for orbitals that are high up in the virtual manifold are typically quite poor. Casida, and Salahub (1998) have suggested that TDDFT results are most reliable if the following two criteria are met: (i) the excitation energy should be significantly smaller than the molecular ionization potential (note that excitations from occupied orbitals below the HOMO are allowed, so this is not a tautological condition) and (ii) promotion(s) should not take place into orbitals having positive KS eigenvalues.

Time-dependent density functional theory is widely used as a reliable method for the prediction of electronic excitation processes and other fast time-dependent phenomena involving the electronic response to an external perturbation. Density functional theory (DFT) accounts for the electron correlation of a system in the ground state by virtue of approximations to the exact but unknown exchange-correlation functional. Thus, TDDFT can also, in principle, incorporate a large fraction of the effects of electron correlation in an excited state within the framework of single-excitation theory. In practice, the accuracy of TDDFT employing the most widely available exchange-correlation functionals, which are approximations to the true functional, and so-called adiabatic approximation has been varied. For vertical excitation energies to low-lying valence excited states, TDDFT with these approximate functionals has proven quite accurate, significantly outperforming the Hartree-Fock based analogues, such as CIS and TDHF. However, for high lying excited states or Rydberg excite states, TDDFT underestimates the excitation energies by up to 1 eV or more.

Ground-state DFT is based on the papers by Hohenberg and Kohn, and by Kohn and Sham. The main results are that the density of system is identical to the

density of an associated noninteracting particle system moving in a local potential  $v_s(\mathbf{r})$  defined by the Kohn-Sham equations (atomic units are used throughout):

$$\left[ -\frac{1}{2}\nabla^2 + v_s[\rho(\mathbf{r})] \right] \phi_i(\mathbf{r}) = \varepsilon_i \phi_i(\mathbf{r}) \quad (39)$$

Here the local potential  $v_s[\rho](\mathbf{r})$  is the so-called Kohn-Sham potential, consisting of the external potential  $v_{\text{ext}}$  (the Coulomb field of the nuclei and external field if present), the Hartree potential  $v_H$ , which is trivially calculated from the density, and the  $x_c$  potential  $v_{xc}$  which is the only unknown part:

$$v_s(\mathbf{r}) = v_{\text{ext}}(\mathbf{r}) + v_H(\mathbf{r}) + v_{xc}(\mathbf{r}) \quad (40)$$

The Kohn-Sham orbitals  $\phi_i$  move in the effective field  $v_s$  which depends upon the electron density  $\rho(\mathbf{r})$ . This density is exactly obtained by summing the squares of the Kohn-Sham orbitals and multiplying by their occupation numbers  $n_i$ .

$$\rho(\mathbf{r}) = \sum_i^{\text{occ}} n_i |\phi_i(\mathbf{r})|^2 \quad (41)$$

As the KS potential  $v_s(\mathbf{r})$  and the density  $\rho(\mathbf{r})$  are inter-dependent, the equations have to be solved in a Self-Consistent Field (SCF) procedure, which means that one iteratively adapts the effective potential  $v_s$  and the density  $\rho$  until the difference in the energy between two subsequent cycles is sufficiently small. In the most straightforward fashion, this can be performed by mixing the density of the previous cycle with a small part of the density in the present cycle. This ‘‘simple damping’’ approach usually converges very slowly, and in practice the Direct Inversion in the Iterative Subspace (DIIS) procedure by Pulay and co-workers, is much to be preferred. In the DIIS approach, not only the result of previous cycle, but the results of all, or many, previous cycles are taken into account, in order to obtain the optimal

guess for the next cycle. If one is close to self-consistency, this procedure converges the SCF equation above.

In order to solve the KS equations an approximation for the exchange-correlation ( $xc$ ) potential  $v_{xc}(r)$  is required and the simplest one is the LDA which is based upon the local density of the system. The GGAs go beyond this and take the local gradient of the density into account as well, allowing for a much improved accuracy in the results for energies and geometries. Many other approximations, for examples those based directly on the KS orbitals, are also available.

The usual ground state DFT scheme enables one to determine the density, and consequently the dipole moment, of a molecule with or without external electric fields. This affords the determination of the static polarizability and hyperpolarizability tensors  $\alpha$ ,  $\beta$  and  $\gamma$  by performing calculations in small electric fields of varying magnitudes and directions. In this so-called finite field (FF) approach, the tensors are then determined from finite difference techniques. The main advantage of this approach is that no programming work is needed. Any standard DFT code will allow the determination of static properties in this manner. However, for the determination of higher order tensors, such as  $\gamma$ , one needs a very well converged solution to the KS equations in order to make reliable predictions, which may be technically hard to achieve and which will certainly lead to a considerable increase in CPU time consuming.

The most fundamental disadvantage of the FF approach, however, is that one has access to static properties only. The frequency-dependent polarizability and hyperpolarizability tensors are not accessible. Excitation energies and oscillator strengths can also not be obtained from the FF calculations. This is an important drawback of the FF approach, as it makes a direct comparison with experimental results impossible. Especially for hyperpolarizabilities, it is known that there are substantial differences between the frequency-dependent and zero frequency results.

If one is interested in the time dependent properties mentioned above, a time dependent theory is required. In the DFT framework, this means that one has to start from the time dependent KS (TDKS) equations as derived by Runge and Gross;

$$i \frac{\partial}{\partial t} \varphi_i(\mathbf{r}, t) = \left[ -\frac{\nabla^2}{2} + v_s(\mathbf{r}, t) \right] \varphi_i(\mathbf{r}, t) \equiv F_s \varphi_i(\mathbf{r}, t) \quad (42)$$

The time dependent KS potential  $v_s(\mathbf{r}, t)$  is subdivided in the same manner as its static counterpart:

$$v_s(\mathbf{r}, t) = v_{\text{ext}}(\mathbf{r}, t) + v_H(\mathbf{r}, t) + v_{\text{xc}}(\mathbf{r}, t) \quad (43)$$

the Hartree potential being explicitly given by:

$$v_H(\mathbf{r}, t) = \int d\mathbf{r}' \frac{\rho(\mathbf{r}', t)}{|\mathbf{r} - \mathbf{r}'|} \quad (44)$$

and the time dependent  $xc$  potential  $v_{\text{xc}}[\rho](\mathbf{r}, t)$  being an unknown functional of the time dependent density  $\rho(\mathbf{r}, t)$  now given by:

$$\rho(\mathbf{r}, t) = \sum_i^{\text{occ}} n_i |\varphi_i(\mathbf{r}, t)|^2 \quad (45)$$

If a certain approximation for the time dependent  $xc$  potential  $v_{\text{xc}}(\mathbf{r}, t)$  has been chosen, the TDKS equations can be solved iteratively to yield the time dependent density of system, which may be exposed to an external time dependent electric field. If one is interested in the effects due to extremely large laser fields, the perturbative expansion of the dipole moment become meaningless, and the TDKS equations have to be solved non-perturbatively. This has until now been performed for atoms, by Ullrich and Gross (Hirata *et al.*, 1999), and more recently also by others, and gives access to such effects as higher harmonic generation (HHG), which are not accessible

in a perturbative approach. The drawback of this is that the calculations are very time consuming, forbidding the treatment of medium-sized molecules. If one restricts oneself to properties which are accessible through perturbative methods, as we will do here, a much more efficient approach is possible, allowing the treatment of large molecules (>100 atoms).

#### 2.4 An Intermediate Neglect of Differential Overlap (INDO) Technique for Spectroscopy

The intermediate neglect of different overlap (INDO) is based on the simplest molecular orbital theory that properly accounts for the two-electron terms required in spectroscopy. INDO/s, where the one-center core integrals are obtained from ionized potentials only, rather than from ionization potentials and electron affinities, is used to calculate the ground-state configuration in terms of molecular orbital coefficients and eigenvalues.

In the molecular orbital model, using the unrestricted Hartree-Fock theory, we are interested in solving the Hartree-Fock equations

$$F^{\alpha}C^{\alpha} = C^{\alpha}E^{\alpha} \quad (46)$$

for an orthogonalized atomic basis set  $\{\phi_{\mu}\}$ , where  $E^{\alpha}$  is the diagonal matrix of eigenvalues for the  $\alpha$  spin molecular orbital,  $C^{\alpha}$  is a square matrix the  $i$  th column of which is the molecular orbital coefficients of the  $i$  th molecular orbital  $\phi_i^{\alpha}$

$$\phi_i^{\alpha} = \sum c_{i\mu}^{\alpha} \phi_{\mu} \quad (47)$$

$F^{\alpha}$  is the Fock or energy matrix

$$F^{\alpha} = H + J - K^{\alpha} \cong H + G^{\alpha} \quad (48)$$

and similar expressions are appropriate for  $\beta$  spin matrices and orbitals. In the above,  $H$  is the one-electron matrix, and  $G^\alpha = J - K^\alpha$  is the two-electron matrix consisting of Coulomb terms minus exchange. The one-electron matrix  $H$  is specified by

$$H_{\mu\mu}^{AA} = U_{\mu\mu}^{AA} - \sum_{B \neq A} Z_B (\phi_\mu | R_B^{-1} | \phi_\mu) \quad (49)$$

$$U_{\mu\mu}^{AA} \cong (\phi_\mu | -\nabla^2/2 - Z_A/R_A | \phi_\mu) \quad (50)$$

$$H_{\mu\nu} = (\phi_\mu | -\nabla^2/2 - \sum_A Z_A/R_A | \phi_\nu) \quad (51)$$

Using the INDO model of Pople, Santry, and Segal we obtained

$$H_{\mu\mu}^{AA} = U_{\mu\mu}^{AA} - \sum_{B \neq A} Z_B \gamma_{AB} \quad (52)$$

$$H_{\mu\nu}^{AA} = 0 \quad (53)$$

$$H_{\mu\nu}^{AB} = (\beta_{A,\mu}^0 + \beta_{A,\mu}^0) \bar{\Delta}_{\mu\nu}/2 \quad (54)$$

$$\gamma_{AB} = \langle \bar{\phi}_\mu \bar{\phi}_\mu | \bar{\phi}_\nu \bar{\phi}_\nu \rangle \cong \int d\tau(1) d\tau(2) \bar{\phi}_\mu(1) \bar{\phi}_\mu(1) r_{12}^{-1} \bar{\phi}_\nu(2) \bar{\phi}_\nu(2) \quad (55)$$

$\{\beta_{A,\mu}^0\}$  are atomic parameters, chosen empirically to give best agreement with experiment. The same parameter is used for either an s or p atomic orbital,  $\beta_{A,s}^0 = \beta_{A,p}^0$ , but a different parameter is chosen for the d orbitals.  $\bar{\Delta}_{\mu\nu}$  is a proportionality constant, usually the orbital overlap calculated treating  $\phi_\mu$  and  $\phi_\nu$  as Slater-type orbitals, or an integral simply related to the overlap.  $\gamma_{AB}$  is the two-electron Coulomb integral calculated treating  $\phi_\mu$  of atom A and  $\phi_\nu$  of atom B as if both were of s symmetry. Such an approximation is required by rotational invariance. Since the orbital exponents of s and p atomic orbitals are chosen the same for a given atom,

$\gamma_{ss}^{AB} = \gamma_{sp}^{AB} = \gamma_{pp}^{AB} = \gamma_{AB}$ . However, the exponent of a d atomic orbital is not the same as an s and p.

In the unrestricted Hartree-Fock formalism that is used here, the G matrix elements are given by

$$G_{\mu\nu}^{\alpha} = \sum_{\sigma\lambda} P_{\sigma\lambda} \langle \mu\nu | \sigma\lambda \rangle - P_{\sigma\lambda}^{\alpha} \langle \mu\sigma | \nu\lambda \rangle \quad (56)$$

with a similar expression for the  $\beta$  electron  $G^{\beta}$  matrix. Here  $P^{\alpha}$  is the first-order  $\alpha$  spin density matrix or, in the orthogonalized set envisioned here, the  $\alpha$  spin charge and bond order matrix

$$P_{\mu\nu}^{\alpha} = \sum_a^{\text{MO}} C_{\mu a}^{\alpha} C_{\nu a}^{\alpha} n_a \quad (57)$$

Where  $n_a = 0$  or  $1$ , the occupancy of  $\phi_a^{\alpha}$ , and P is the total first-order density

$$P = P^{\alpha} + P^{\beta} \quad (58)$$

Under this formalism, for a closed-shell system

$$P^{\alpha} = P^{\beta} = 1/2P \quad (59)$$

Invoking the INDO model simplifies eq. (55) greatly:

$$G_{\mu\mu}^{\alpha} = \sum_{\sigma\lambda}^A [P_{\sigma\pi} \langle \mu\mu | \sigma\lambda \rangle - P_{\sigma\lambda}^{\alpha} \langle \mu\sigma | \nu\lambda \rangle] + \sum_{\sigma \notin A} P_{\sigma\sigma} \gamma_{\sigma\mu}; \mu \in A \quad (60)$$

$$G_{\mu\nu}^{\alpha} = \sum_{\sigma\lambda}^A [P_{\sigma\pi} \langle \mu\nu | \sigma\lambda \rangle - 1/2 P_{\sigma\lambda}^{\alpha} \langle \mu\sigma | \nu\lambda \rangle]; \mu, \nu \in A \quad (61)$$

$$G_{\mu\nu}^{\alpha} = -P_{\mu\nu}^{\alpha} \gamma_{\mu\nu}^{-}; \mu \in A, \nu \in B, A \neq B \quad (62)$$

with similar expressions for  $G^{\beta}$ .

The two-electron two-center Coulomb integrals  $\gamma_{\mu\nu}^{-}$  are required for eq. (55) and for eq (62). For spectroscopy, however, a great deal of experience has been gained by using Coulomb integrals chosen semiempirically.

The two-electron Coulomb integrals are evaluated from a modified Mataga-Nishimoto formula

$$\gamma_{AB} = \frac{f_{\gamma}}{2f_{\gamma}/(\gamma_{AA} + \gamma_{BB}) + R_{AB}} \quad (63)$$

where  $R_{AB}$  is the distance between the two centers in Bohr radii,  $\gamma_{AA}$  is obtained from

$$\gamma_{AA} = F^0(AA) = I_A - A_A \quad (64)$$

as the difference between the ionization potential and electron affinity of an s, p or d electron. The formula is modified by the introduction of the parameter  $f_{\gamma}$  as suggested by Weiss, which we set equal to 1.2. This value has been extensively used and has proven very successful in the calculation of spectra of conjugated systems. Since values of  $I_A - A_A$  are similar for s and p orbitals, we make the approximation that  $\gamma_{\overline{AA}} = \gamma_{\overline{ss}} = \gamma_{\overline{sp}} = \gamma_{\overline{pp}} = F^0(AA)$  as before.

The INDO spectroscopic model has been parametrized by comparison of experimental spectra with calculated results obtained after extensive studies with singly excited configurations in a configuration interaction treatment. This basic model has been extended to include members of the first transition series (Zerner *et al.*, 1979), and has been applied to rather large systems in which the role of higher

excited configurations (double excited, triple excited, etc., as opposed to higher energy configurations) is uncertain. Although use of semiempirical two-electron integrals ( $\gamma$ ) and empirical resonance integrals ( $\beta$ ) might be expected to include the effect of higher energy excitations in an average way, no such parametrization can be expected to compensate for multiple excited configurations if those configurations are actually in the spectroscopic region of interest, or if a given double excitation heavily mixes with the reference configuration in the mathematical description of the ground state. In these cases, certain higher excitations *must* be included explicitly. Preliminary investigations seem to indicate that higher excitation excitations can be included in the model without extensive reparamitization. Nevertheless, some systems cannot be well described without these higher excitations, even with the given semiempirical parametrization. The frequency of these more difficult to handle systems is expected to increase with increasing molecular size, and the effect will be aggravated by the presence of a transition metal. The presence of two like transition metals in most complexes will require consideration of higher excitations in the description of both ground and excited states.

In addition to the model being grounded on singly excited configuration interaction, which is not a severe limitation in most cases, the model is founded on basis set of valence-type orbitals. Since Rydberg-type orbitals are not included in the atom basis, Rydberg-type states cannot be described. This is as severe a limitation on small molecules where the Rydberg states are among the lowest lying excited states as exclusion of higher excitations is for very large systems, where they are low lying. For most molecules Rydberg states might be expected to make their appearance 2-3 eV below the first ionization potential. In this region of the spectrum the density of states is quite high. The appearance of broad absorptions usually found within 2-3 eV of the first ionization potential for large systems may have as much to do with the presence of many electronic origins as it has to do with the vibrational broadening usually considered.

## The Theory of Photophysical Processes

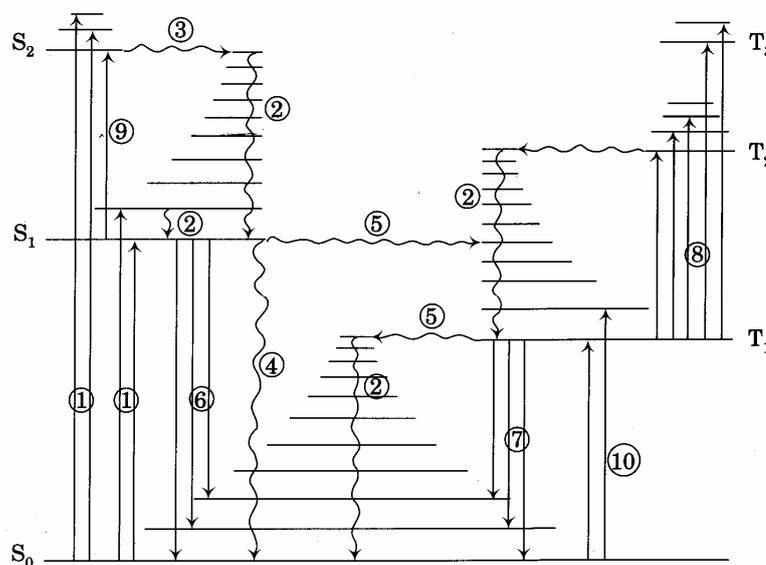
### 1. Photophysical Process

If a substance is irradiated with electromagnetic radiation, the energy of the incident photons may be transferred to the atoms or molecules raising them from the ground state to an excited state. This process, known as *absorption*, is accompanied by attenuation of the incident radiation at a particular frequency and can occur only when the energy difference between the two levels is exactly matched by the energy of the photons. The frequency of the radiation is given by

$$E_2 - E_1 = \Delta E = h\nu \quad (1)$$

where  $E_1$  and  $E_2$  are the energies of the two levels and  $\Delta E$  is the difference between them. The absorbed energy is rapidly lost to the surrounding by collisions allowing the system to revert or relax to the ground state. Sometimes the energy is not dissipated in this way but is re-emitted a few million seconds later—a process known as *fluorescence*.

A useful way of representing the energies of the electronic states of molecule is a **Jablonski diagram**, such as in Fig. A3. The vertical scale represents potential energy. The horizontal scale has no particular significance; it allows us to separate to singlet and triplet state **manifolds** (sets of energy levels) so that the excited states are more clearly distinguished. Superimposed on each electronic state is a set of vibrational energy levels. For the sake of clarity, a set of rotational energy levels superimposed on each vibrational level is not shown.



**Appendix Figure A3** Generalized Jablonski diagram.

There are two kinds of photophysical processes indicated in Fig. A3. Those interconversions denoted by straight lines are **radiative processes**, which occur through the absorption or emission of light. Those indicated by wavy lines are **nonradiative processes**, which occur without light being absorbed or emitted. The numbers on the lines are keys to the following definitions:

**1. Absorption of Light.** A ground state molecule ( $S_0$ ) may absorb a photon of light, thus becoming converted to an excited state. The most likely transitions are  $S_0 \rightarrow S_1$  or  $S_0 \rightarrow S_2$ , although  $S_0$  to higher excited singlet state transitions are also possible.

**2. Vibrational Relaxation.** The absorption from  $S_0$  to  $S_n$  involves an energy change from the 0<sup>th</sup> vibrational level of  $S_0$  to any vibrational level of the excited state. However, the  $v'' = 0$  is also the vibrational level is the level most populated at room temperature for the ground electronic state of molecule, and  $v' = 0$  is also the vibrational level of the excited electronic state that is most likely to be populated at equilibrium. Unless the molecule dissociates before equilibrium can be obtained, there

is a very rapid process (with a rate constant of about  $10^{12} \text{ sec}^{-1}$ ) that relaxed the higher vibrational level of the excited state to its 0<sup>th</sup> vibrational level in condensed phases (i.e., solids or liquids).

**3. Internal Conversion** is a non radiative process that converts a higher electronic state into a lower state of the same multiplicity (a higher singlet state into a lower singlet state or a higher triplet state into a lower triplet state). The name arises because the process occurs internally, the line for internal conversion is a horizontal one. This means that the the 0<sup>th</sup> vibrational level of  $S_2$  is converted into a vibrationally excited  $S_1$  can then relax to its 0<sup>th</sup> vibrational level. The rate constants for internal conversion are fast ( $> 10^{10} \text{ sec}^{-1}$ ), especially when the two states are close in energy.

**4. Radiationless Decay** is a process by which electronically excited states are returned to ground states (typically from  $S_1$  to  $S_0$ ) without the emission of light. Radiationless decay often has a slower rate constant (ca.  $<10^6 \text{ sec}^{-1}$ ) than other forms of internal conversion because the energy gap between  $S_1$  and  $S_0$  is usually greater than that between  $S_2$  and  $S_1$  or other pairs of excited states.

**5. Intersystem Crossing**, the conversion of a singlet state into a triplet state (or vice versa), requires a spin flip of an electron. The probability of intersystem crossing depends, among other things, on the energy gap between the singlet and triplet states, so values of  $k_{isc}$  vary from  $10^6$  to  $10^{10} \text{ sec}^{-1}$ . The  $T_n$  states is lower in energy than the corresponding  $S_n$  state because of the lower electron repulsion for unpaired electrons.

Processes 2-5 are nonradiative processes. The following are radiative processes.

**6. Fluorescence** is the emission of light from an excited state to a ground state with the same multiplicity. Usually the emission is  $S_1 \rightarrow S_0$ , and generalization to that effect is known as Kasha's rule. However, **anomalous fluorescence** ( $S_2 \rightarrow S_0$ ) occurs in some compound, among them azulene, thiocarbonyl compounds, and some gaseous polyenes.

**7. Phosphorescence** is the emission of light from an excited state to a ground state with different multiplicity (usually from a triplet excited state to singlet ground state). This process involves both an electronic state change and a spin flip so, like  $S_0 \rightarrow T_n$  absorption, phosphorescence is a spin-forbidden process.

**8. Triplet-Triplet Absorption.** A molecule in a triplet excited state may absorb a photon to give a higher triplet state, so a UV-Vis spectrum may be obtained, and the time-dependence of the excited state decay may be monitored. Flash spectroscopy is triplet-triplet absorption spectroscopy that can be an important technique for detecting triplet excited states.

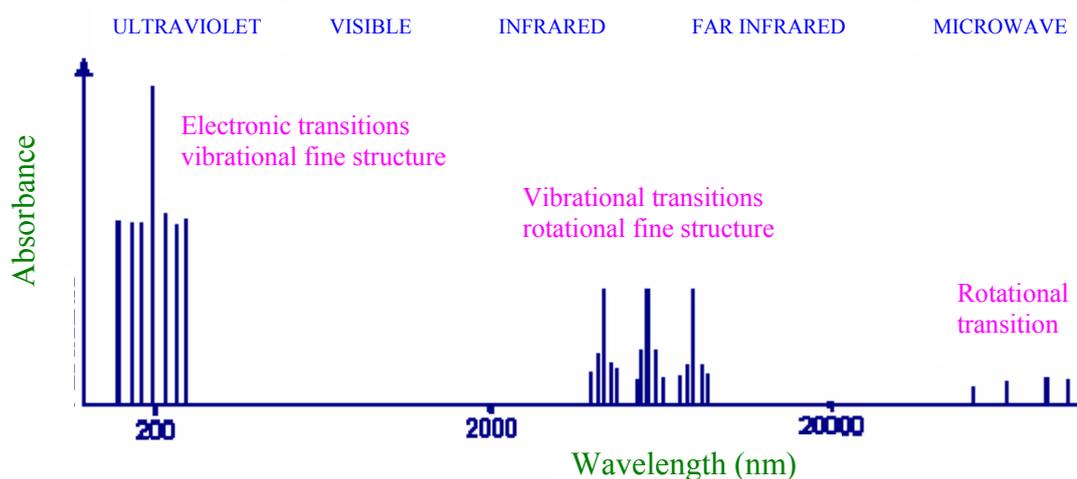
**9. Singlet-Singlet Absorption.** Because triplet states may persist longer than singlet states, the only excitation of excited states possible during much of the development of photochemistry was triplet-triplet absorption. With the advent of picosecond and femtosecond spectroscopy, it has become possible also to measure transitions from one excited singlet state to another, higher energy excited singlet state.

**10. Singlet-Triplet Absorption,** like phosphorescence, is a spin-forbidden process, so ordinarily  $S_0 \rightarrow T_n$  transitions are not observed in UV-Vis spectroscopy. However, these transitions can be seen under certain conditions.

For ultraviolet and visible wavelengths, one should expect that the absorption spectrum of a molecule (i.e., a plot of its degree of absorption against the wavelength of the incident radiation) should show a few very sharp lines. Each line should occur at a wavelength where the energy of an incident photon exactly matches the energy required to excite an electronic transition as shown in Fig. A4.

In practice it is found that the ultraviolet and visible spectrum of most molecules consists of a few humps rather than sharp lines. These humps show that the molecule is absorbing radiation over a band of wavelengths. One reason for this band, rather than line absorption is that an electronic level transition is usually accompanied by a simultaneous change between the more numerous vibrational levels. Thus, a

photon with a little too much or too little energy to be accepted by the molecule for a ‘pure’ electronic transition can be utilized for a transition between one of the vibrational levels associated with the lower electronic state to one of the vibrational levels of a higher electronic state.



**Appendix Figure A4** Idealized absorption spectrum.

If the difference in electronic energy is ‘E’ and the difference in vibrational energy is ‘e’, then photons with energies of E, E+e, E+2e, E-e, E-2e, etc. will be absorbed.

Furthermore, each of the many vibrational levels associated with the electronic states also has a large number of rotational levels associated with it. Thus a transition can consist of a large electronic component, a smaller vibrational element and an even smaller rotational change. The rotational contribution to the transition has the effect of filling in the gaps in the vibrational fine structure. In addition, when molecules are closely packed together as they normally are in solution, they exert influences on each other which slightly disturb the already numerous, and almost infinite energy levels and blur the sharp spectral lines into bands.

## 2. Fluorescence and Phosphorescence

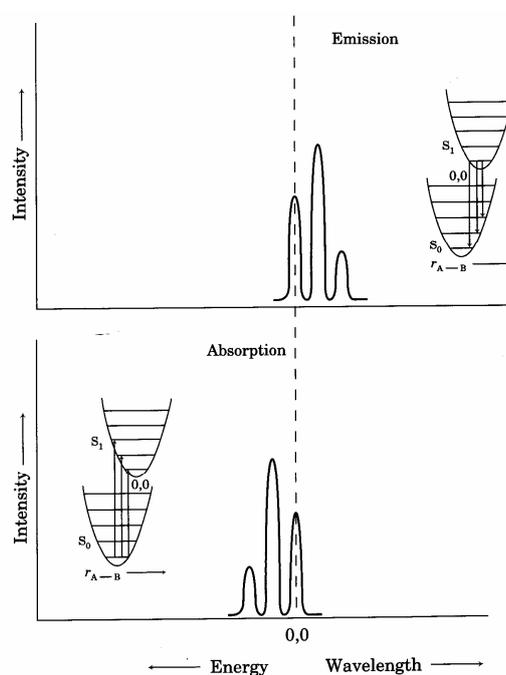
For most organic molecules, fluorescence is the spontaneous emission of light from the  $v' = 0$  vibrational level of the first excited singlet state to some vibrational level ( $v'' = 0, 1, 2, \dots$ ) of the (singlet) ground electronic state. As indicated in the top portion of Fig. A5, the energy of the photon emitted in the  $v' = 0$  to  $v'' = 0$  fluorescence is the same as the energy of the photon absorbed in the  $v'' = 0$  to  $v' = 0$  transition if the geometry of the photoexcited molecule is nearly the same as that of the ground state molecule. However, all other fluorescence lines are at longer wavelengths (lower energy). Thus, the fluorescence and emission spectra should overlap at the 0,0 transition, providing confirmation of the 0,0 energy of the electronically excited state. In molecules such as anthracene, the  $\sigma$  bonding provides a molecular framework for the planar  $\pi$  system, and the  $\pi$  bonding results from population of many bonding MOs. Therefore promotion of one electron to an antibonding MO may not seriously distort the molecular geometry. In such cases there can be a nearly mirror image relationship between the absorption and fluorescence spectra of organic molecules, particularly when the spectra are plotted as intensity versus energy ( $\text{cm}^{-1}$ ) instead of wavelength. The similarity arises because the factors that make some  $v'' = 0$  to  $v' = x$  transitions more probable than others also make some  $v' = 0$  to  $v'' = x$  emissions more probable.

For some molecules, however, the geometry of the ground and excited states may be very different. As a result, there may be a large difference between  $\lambda_{\text{max}}$  for absorption and  $\lambda_{\text{max}}$  for emission, and the 0,0 transition may be weak or not present.

A similar relationship between geometry and a mirror image appearance of singlet-triplet absorption and phosphorescence is expected, but this is often difficult to determine experimentally. In fluid solution phosphorescence is usually reduced by diffusion-limited bimolecular interaction of the excited triplet compound and one or more ground state species. However, phosphorescence can often be observed by

irradiating the compound in environments in which diffusion is quite slow, such as in an organic glass at liquid nitrogen temperature.

Singlet-triplet absorption is ordinary difficult to detect because the transition is spin-forbidden. Heavy atom solvents or oxygen perturbation have been used to induce singlet-triplet absorption, but absorptions observed in this manner are weak, and it is important to establish that the observed absorption is not due to artifacts resulting from the solvent or additive.



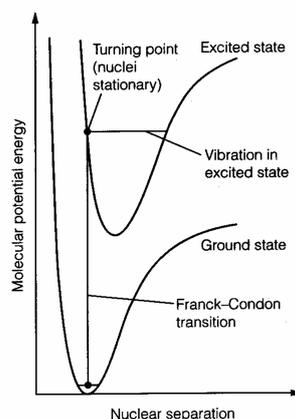
**Appendix Figure A5** Schematic representations of the origins of UV-Vis absorption (bottom) and fluorescence (top) spectra.

### 3. Franck-Condon principle

The different electronic states of a molecule are often associated with different shapes of molecule because the different electron distribution around the molecule changes the electrostatic Coulombic forces that maintain the nuclei in specific relative positions. Since nuclei are considerably more massive than electrons, the **Franck-Condon principle** states that:

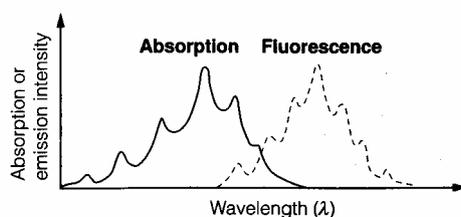
*an electronic transition takes place sufficiently rapidly that the nuclei do not change their internuclear positions during the transition.*

Consequently, when energy is absorbed in an electronic transition, the nuclei suddenly find themselves in a new force field and at positions which are not in equilibrium for the new electronic state. This is shown schematically in Fig. A6, in which an electronic absorption from the ground state appears as a vertical line because of the Franck-Condon principle. The internuclear separation of the ground state becomes a **turning point**, the extent of maximum displacement, in a vibration of the excited state.



**Appendix Figure A6** Illustration of the Franck-Condon principle for vertical electronic transitions.

The vertical transition has the greatest **transition probability** but transitions to nearby **vibrational levels** also occur with lower intensity. Therefore, instead of an electronic absorption occurring at a single, sharp line, electronic absorption consists of many lines each corresponding to the stimulation of different vibrations in the upper state. This vibrational structure (or progression) of an electronic transition can be solved for small molecules in the gas-phase, but in a liquid or solid **collision broadening** of the transitions cause the lines to merge together and the electronic absorption spectrum is often a broad band with limited structure (Fig. A7). The Franck-Condon principle also applies to downward transitions and accounts for the vibrational structure of a **fluorescence** spectrum.



**Appendix Figure 7** Relationship between the broad electronic absorption and fluorescence bands of liquids and solids.